

- Final on Saturday July 30th at SOLIS 104; 7 to 10 pm

Numerically solving DEs

Consider DE $\frac{d}{dt}x = f(t, x)$ $x: I \rightarrow \mathbb{R}^n$
 $f: I \rightarrow \mathbb{R}^n$

Say f is cont. diff. and we're given
 $x(0) = x_0$.

Def: A discretization of the IVP $\begin{cases} \frac{d}{dt}x = f(t, x) \\ x(0) = x_0 \end{cases}$

is a sequence of approximate problems that can be solved (e.g., on a computer). It is consistent if the approx. error goes to 0 as we go further in this sequence.

[The Forward Euler Method]

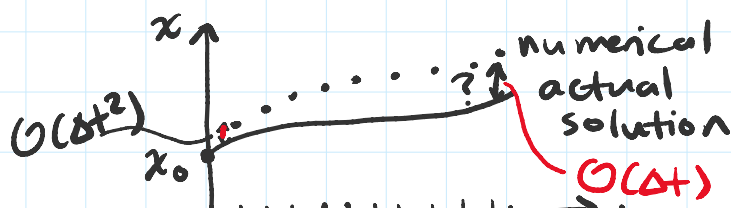
$$\begin{aligned} x(t+\Delta t) &= x(t) + \Delta t x'(t) + \mathcal{O}(\Delta t^2) \\ &= x(t) + \Delta t f(t, x(t)) + \mathcal{O}(\Delta t^2) \end{aligned}$$

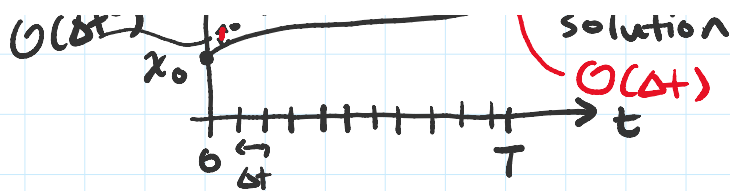
The forward Euler method says ignore \uparrow

$$t_n = n \Delta t \quad n=1, 2, 3, \dots$$

$$x_n = x(n \Delta t)$$

$$\rightarrow x_{n+1} = x_n + \Delta t f(t_n, x_n), \quad x_0 = x(0) = x_0$$





Fix $T > 0$, want to know the solution on $[0, T]$

Number of steps $N = \frac{T}{\Delta t}$ as $\Delta t \rightarrow 0$, T fixed,
 $N \rightarrow \infty$

Total error

= (number of timesteps) \times (error per step)

= $\frac{T}{\Delta t} \times \mathcal{O}(\Delta t^2) = \mathcal{O}(\Delta t)$.

first-order because global error Δt^1 .

$\rightarrow 0$ as $\Delta t \rightarrow 0$

Def: A discretization is a reduction

of a problem w/ an infinite-dimensional candidate solution space to one with a finite-dim. candidate solution space. It is consistent if the approx error goes to zero as the finite-dimension goes to ∞ .

$\frac{d}{dt} x = f(t, x)$
 $x(0) = x_0$

Search for $x \in C^2([0, T])$.

Infinite-dim. vector space.

approximation space

$\mathbb{R}^N \leftarrow$ number of timesteps

Another 1st-order method: Backward-Euler (BE)

$$x_{n+1} = x_n + \Delta t f(t_{n+1}, x_{n+1})$$

algebraic equation to solve for x_{n+1} .

⇒ More expensive to solve than FE.

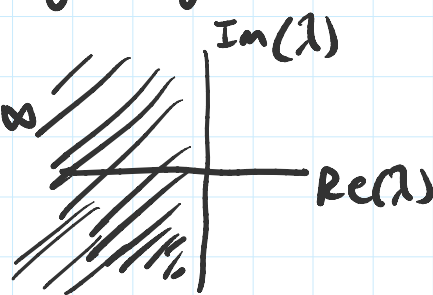
Why would I use this? Stability
(Linear Stability Analysis)

consider $\frac{d}{dt} y = \lambda y$

$$y(0) = y_0$$

True solution $y(t) = y_0 e^{\lambda t}$

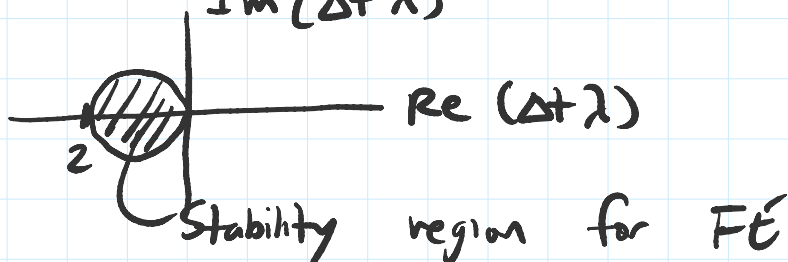
If $\text{Re}(\lambda) < 0$, $y(t) \rightarrow 0$ as $t \rightarrow \infty$



FE:

$$\begin{aligned} y_{n+1} &= y_n + \Delta t \lambda y_n \\ &= (1 + \Delta t \lambda) y_n \\ &= (1 + \Delta t \lambda)^{n+1} y_0 \end{aligned}$$

$|1 + \Delta t \lambda| < 1$ for $y_n \rightarrow 0$ as $n \rightarrow \infty$



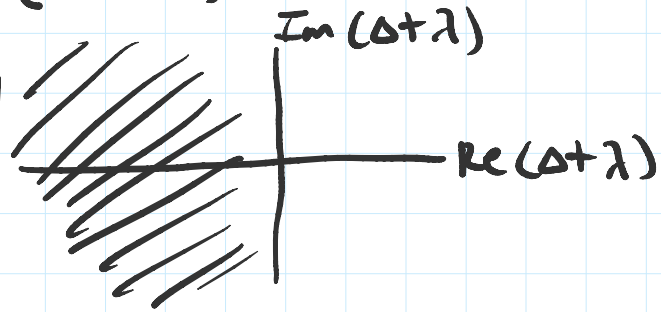
Assume $\lambda \in \mathbb{R}$, $\lambda < 0$, $0 < \Delta t < -\frac{2}{\lambda}$.

↑
restricts timestep

BE: $y_{n+1} = y_n + \Delta t \lambda y_{n+1}$

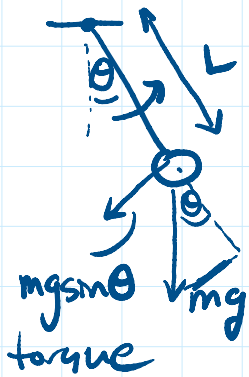
$$y_{n+1} = \frac{y_n}{1 - \Delta t \lambda} = \frac{y_0}{(1 - \Delta t \lambda)^{n+1}}$$

For stability, $|1 - \Delta t \lambda| > 1$



BE is unconditionally stable

Let's consider an ex:



Newton's 2nd law for rotational motion

$$\text{torque} = (\text{moment of inertia}) \times (\text{angular accel.})$$

$$-mgL \sin \theta = mL^2 \frac{d^2 \theta}{dt^2}$$

$$\Rightarrow \frac{d^2 \theta}{dt^2} = -\frac{g}{L} \sin \theta \quad \left(\begin{array}{l} \text{linearize} \\ \frac{d^2}{dt^2} \theta = -\frac{g}{L} \theta + \mathcal{O}(\theta^3) \end{array} \right)$$

Note: energy $E = \frac{1}{2} \left(\frac{d\theta}{dt} \right)^2 - \frac{g}{L} \cos \theta$ is conserved

$$\frac{dE}{dt} = \frac{d\theta}{dt} \frac{d^2 \theta}{dt^2} + \frac{g}{L} \sin \theta \frac{d\theta}{dt}$$

$$= \frac{d\theta}{dt} \left(\frac{d^2 \theta}{dt^2} + \frac{g}{L} \sin \theta \right) = 0.$$

$$= \frac{dv}{dt} \left(\underbrace{\frac{dv}{dt} + \frac{v}{L} \sin \theta}_{=0} \right) = 0.$$

Let's solve w/ FE

$$v = \frac{d\theta}{dt}$$

$$\frac{dv}{dt} = -\frac{g}{L} \sin \theta \quad g=l=L$$

FE reads

$$\rightarrow \theta_{n+1} = \theta_n + \Delta t v_n$$

$$\rightarrow v_{n+1} = v_n - \Delta t \sin \theta_n$$

BE

$$\theta_{n+1} = \theta_n + \Delta t v_{n+1} \quad \leftarrow$$

$$v_{n+1} = v_n - \Delta t \sin \theta_{n+1} \quad \leftarrow$$

Symplectic Euler (SE)

$$\text{FE in } \theta : \quad \theta_{n+1} = \theta_n + \Delta t v_n \quad \leftarrow$$

$$\text{BE in } v : \quad v_{n+1} = v_n - \Delta t \sin \theta_{n+1} \quad \leftarrow$$

SE doesn't preserve E exactly, but it oscillates about the constant energy surface.

This is the idea "structure-preserving discretizations"

SE preserves the symplectic structure exactly.
Symplectic structure is more important than energy conservation.

[Partial Differential Equations]

PDEs generalize ODEs to allow derivatives in more than 1 variable. Describe fields evolving in space time:

Maxwell's equations - electromagnetic fields

Schrödinger's equation - quantum mechanics

Einstein's field eqns - curvature of spacetime

Navier-Stokes - fluid dynamics

Wave Equation: describes oscillations of some physical quantity propagating through space and time

ex/ light

ex/ gravitational waves

Waves on a String

Waves on a String

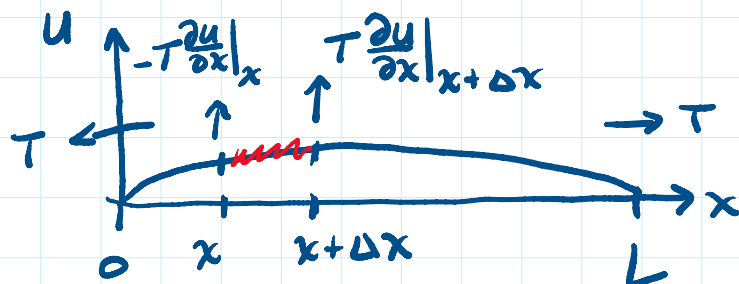
Consider a string clamped at two ends

$$x=0, x=L$$



Suppose the string has linear density ρ (mass/unit length) and tension T

Let $u(t, x)$ denote the displacement of the string at time t and position x .



"Boundary Conditions"

$$u(t, 0) = 0$$

$$u(t, L) = 0$$

$$F = ma$$

$$T \left(\frac{\partial u}{\partial x} \Big|_{x+\Delta x} - \frac{\partial u}{\partial x} \Big|_x \right) = \overset{m}{\rho \Delta x} \frac{\partial^2}{\partial t^2} u(t, x) + \mathcal{O}(\Delta x^2)$$

Divide by Δx and let $\Delta x \rightarrow 0$

$$T \left(\frac{\frac{\partial u}{\partial x} \Big|_{x+\Delta x} - \frac{\partial u}{\partial x} \Big|_x}{\Delta x} \right) = \rho \frac{\partial^2}{\partial t^2} u + \underbrace{\mathcal{O}(\Delta x)}_0$$

$$\frac{\partial^2 u}{\partial x^2}$$

$$\text{Wave Equation: } -\partial^2 u = \rho \partial^2 u$$

Wave Equation $T \frac{\partial^2 u}{\partial x^2} = \rho \frac{\partial^2 u}{\partial t^2}$

$$\frac{\partial^2 u}{\partial x^2} = \frac{1}{c^2} \frac{\partial^2 u}{\partial t^2} \quad c = \sqrt{T/\rho}$$

↑ speed

BCs $\begin{cases} u(t, 0) = 0 \\ u(t, L) = 0 \end{cases}$

ICs $\begin{cases} u(0, x) = u_0(x) \\ \frac{\partial u}{\partial t}(0, x) = v_0(x) \end{cases}$

Solve this using separation of variables

$$u(t, x) = X(x)T(t) \quad X(0) = 0 = X(L)$$

$$\frac{\partial^2 u}{\partial x^2} = X''(x)T(t) \quad \frac{\partial^2 u}{\partial t^2} = X(x)T''(t)$$

$$X''(x)T(t) = \frac{1}{c^2} X(x)T''(t)$$

$$\frac{X''(x)}{X(x)} = \sigma = \frac{1}{c^2} \frac{T''(t)}{T(t)}$$

↑ const.

$$\sigma > 0$$

$$X''(x) = \sigma X(x)$$

~~$$X(x) = c_1 e^{\sqrt{\sigma}x} + c_2 e^{-\sqrt{\sigma}x}$$~~

$$\sigma = 0$$

~~$$v''(x) = 0$$~~

~~$$v''(t) = 0$$~~

$$\sigma = 0 \quad X''(x) = 0 \Rightarrow \underline{X(x) = C_1 x + C_2}$$

$$\sigma < 0 \quad \alpha = -\sigma > 0$$

$$X''(x) = -\alpha X(x)$$

↑ harmonic oscillator

$$X(x) = C_1 \sin(\sqrt{\alpha} x) + C_2 \cos(\sqrt{\alpha} x)$$

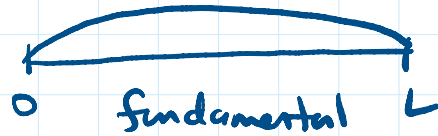
$$X(0) = 0 \quad \checkmark$$

$$X(L) = 0$$

When does $\sqrt{\alpha} L$ satisfy $\sin(\sqrt{\alpha} L) = 0$

$$\sqrt{\alpha} L = k\pi \quad k = 1, 2, 3, \dots$$

$$X_k(x) = \underline{C_k} \sin\left(\frac{k\pi}{L} x\right) \quad k=1$$



$$T''(t) = -\alpha C^2 T(t)$$

$$\Rightarrow T_k(t) = \beta_k \cos\left(\frac{Ck\pi}{L} t\right) + \gamma_k \sin\left(\frac{Ck\pi}{L} t\right)$$

general solution

linear superposition

$$u(t, x) = \sum_{k=1}^{\infty} X_k(x) T_k(t)$$

$$= \sum_{k=1}^{\infty} \sin\left(\frac{k\pi}{L} x\right) \left(\beta_k \cos\left(\frac{Ck\pi}{L} t\right) + \gamma_k \sin\left(\frac{Ck\pi}{L} t\right) \right)$$

$$u_0(x) = u(0, x) = \sum_{k=1}^{\infty} \beta_k \sin\left(\frac{k\pi}{L} x\right)$$

$$v_0(x) = \frac{\partial u}{\partial t}(0, x) = \sum_{k=1}^{\infty} \frac{ck\pi}{L} \gamma_k \sin\left(\frac{k\pi}{L} x\right)$$

$$\beta_k = \frac{2}{L} \int_0^L u_0(x) \sin\left(\frac{k\pi}{L} x\right) dx$$

$$\gamma_k = \frac{2}{ck\pi} \int_0^L v_0(x) \sin\left(\frac{k\pi}{L} x\right) dx$$

Fourier sine series.

$$\text{frequency} \quad \frac{ck\pi}{L} = \sqrt{\frac{T}{\rho}} \frac{k\pi}{L}$$